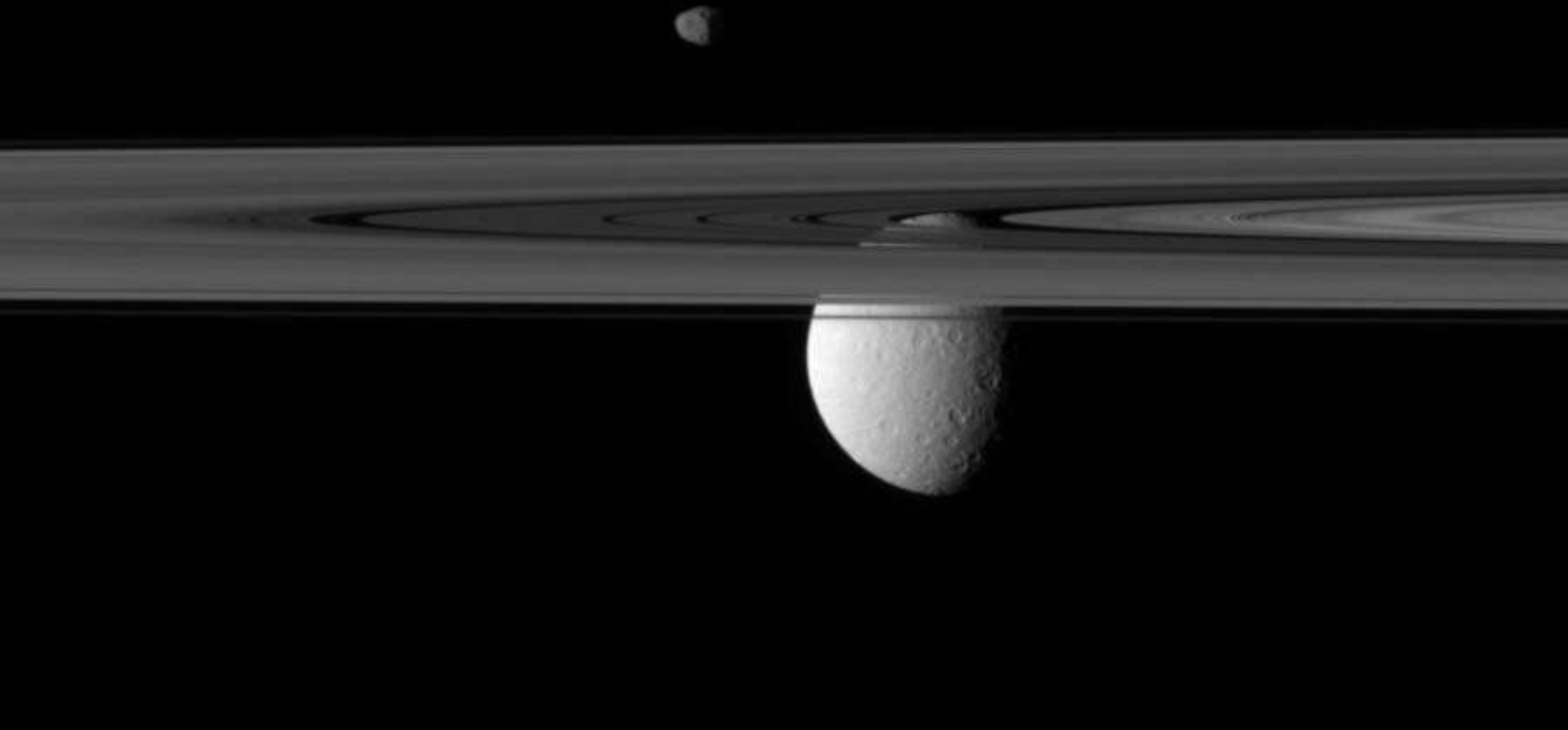


From large velocities to tiny accelerations

- the successful Doppler detection of exo-planets

gordon walker



spectrum display

complete UBC Isocon
digital camera system
1970

ADC
exposure time

Interdata computer
4K memory
+ tape drive

Isocon
controls

Isocon

monitors

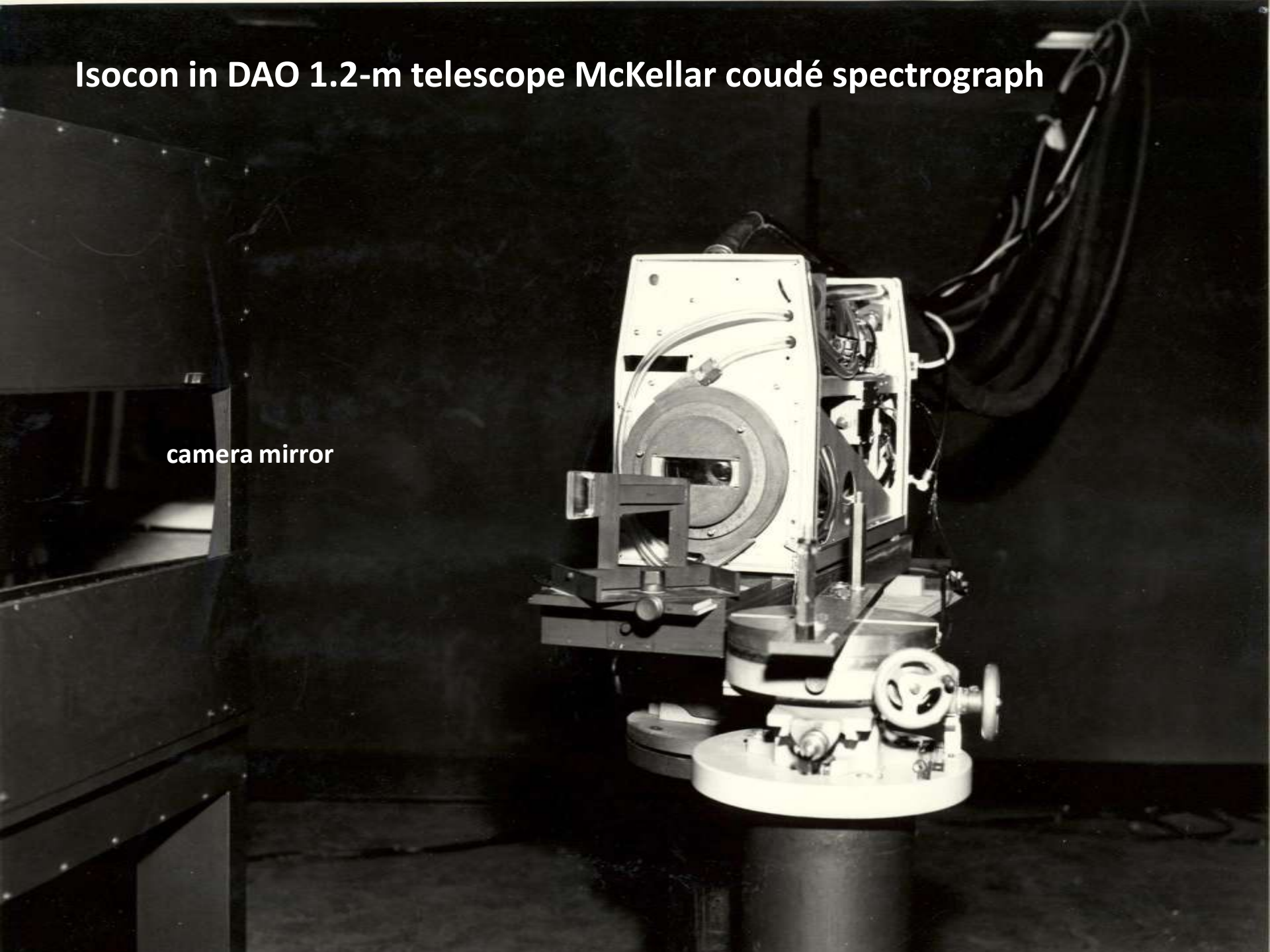
cooling system

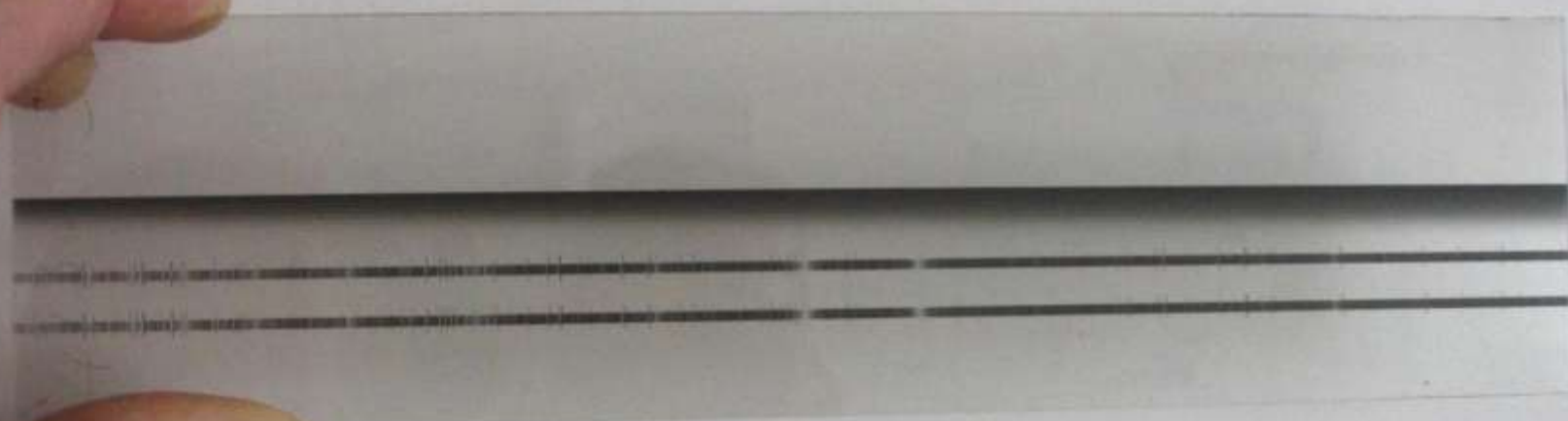
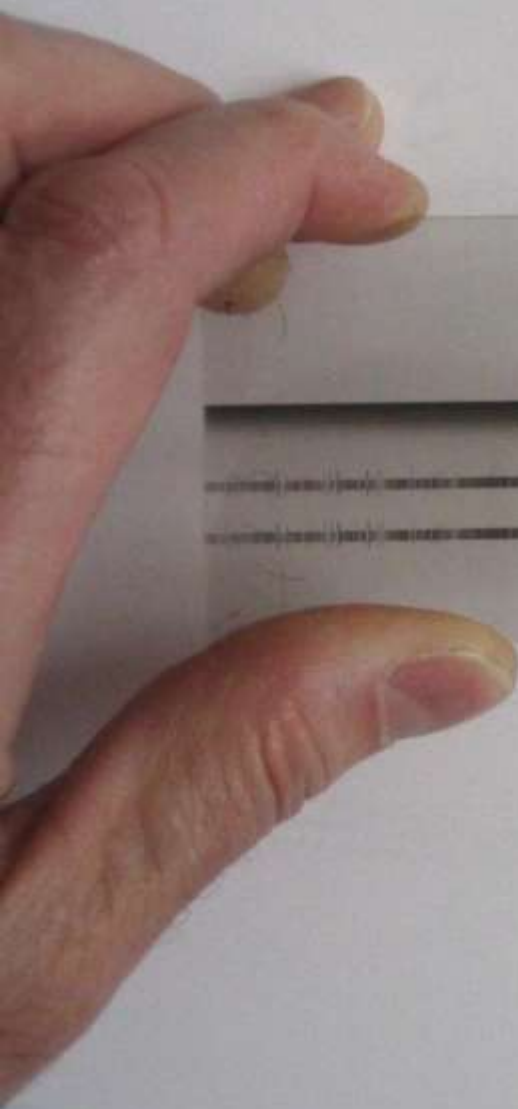


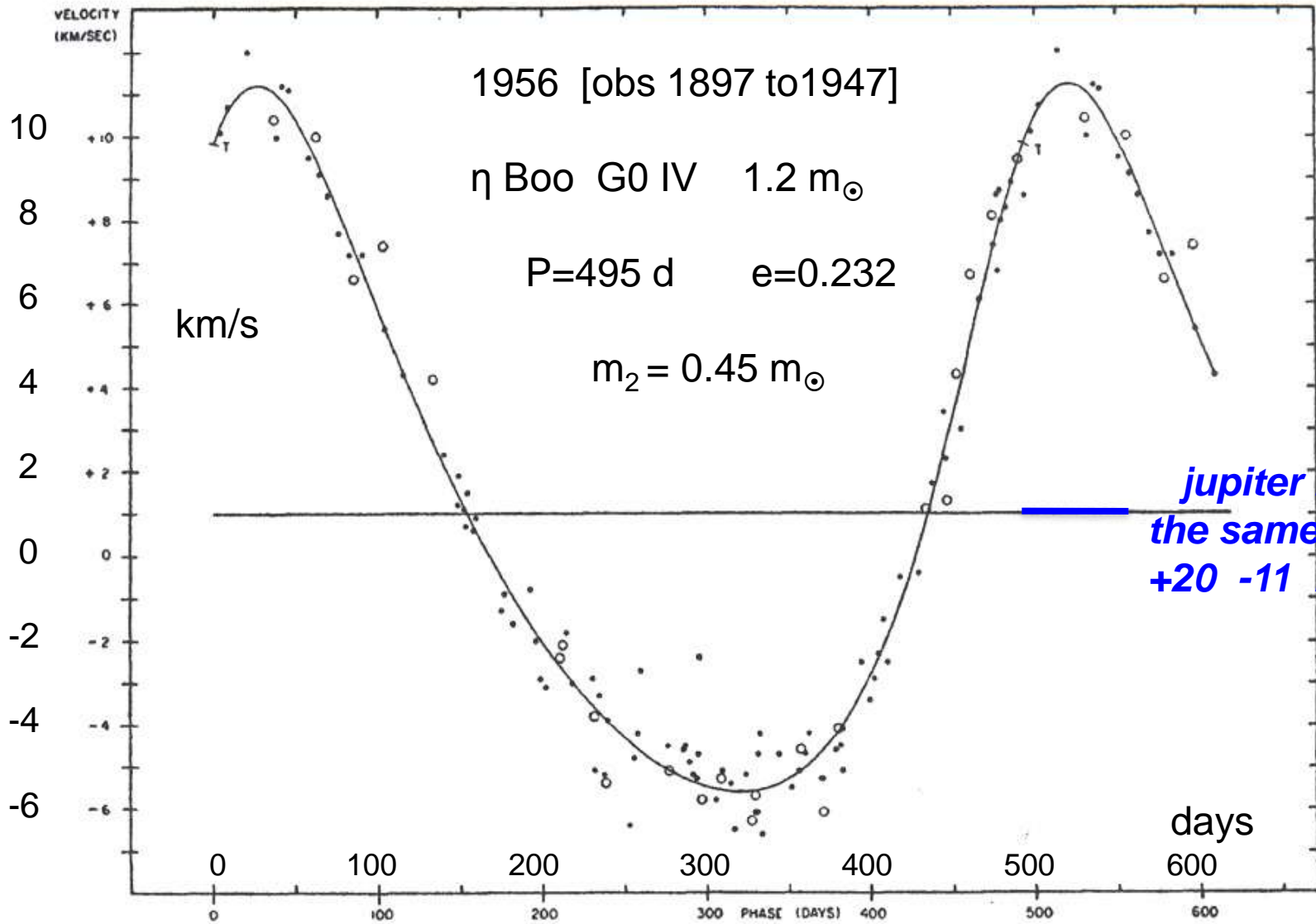
C_2H_5OH

Isocon in DAO 1.2-m telescope McKellar coudé spectrograph

camera mirror



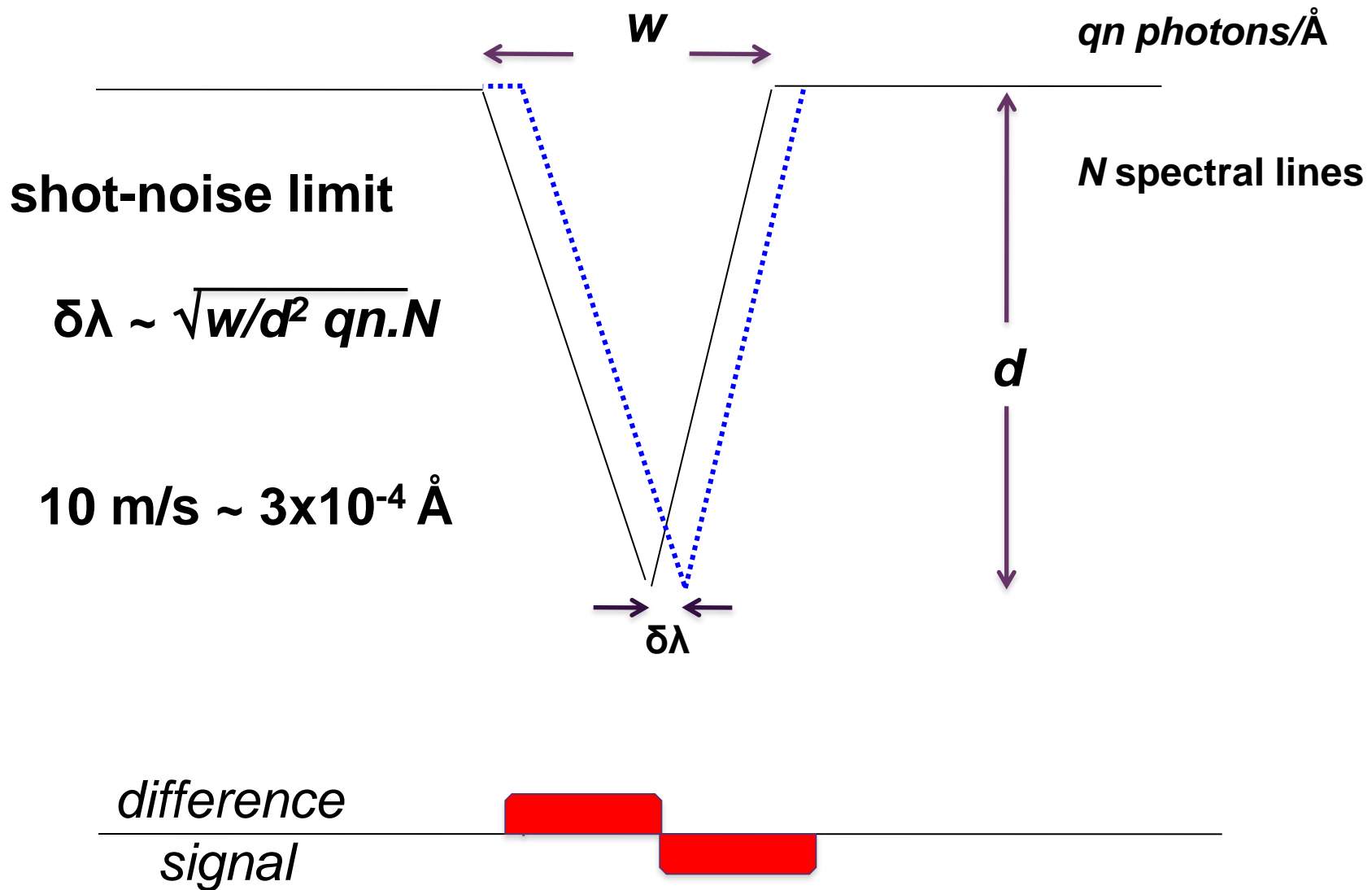




~ 10 km/s



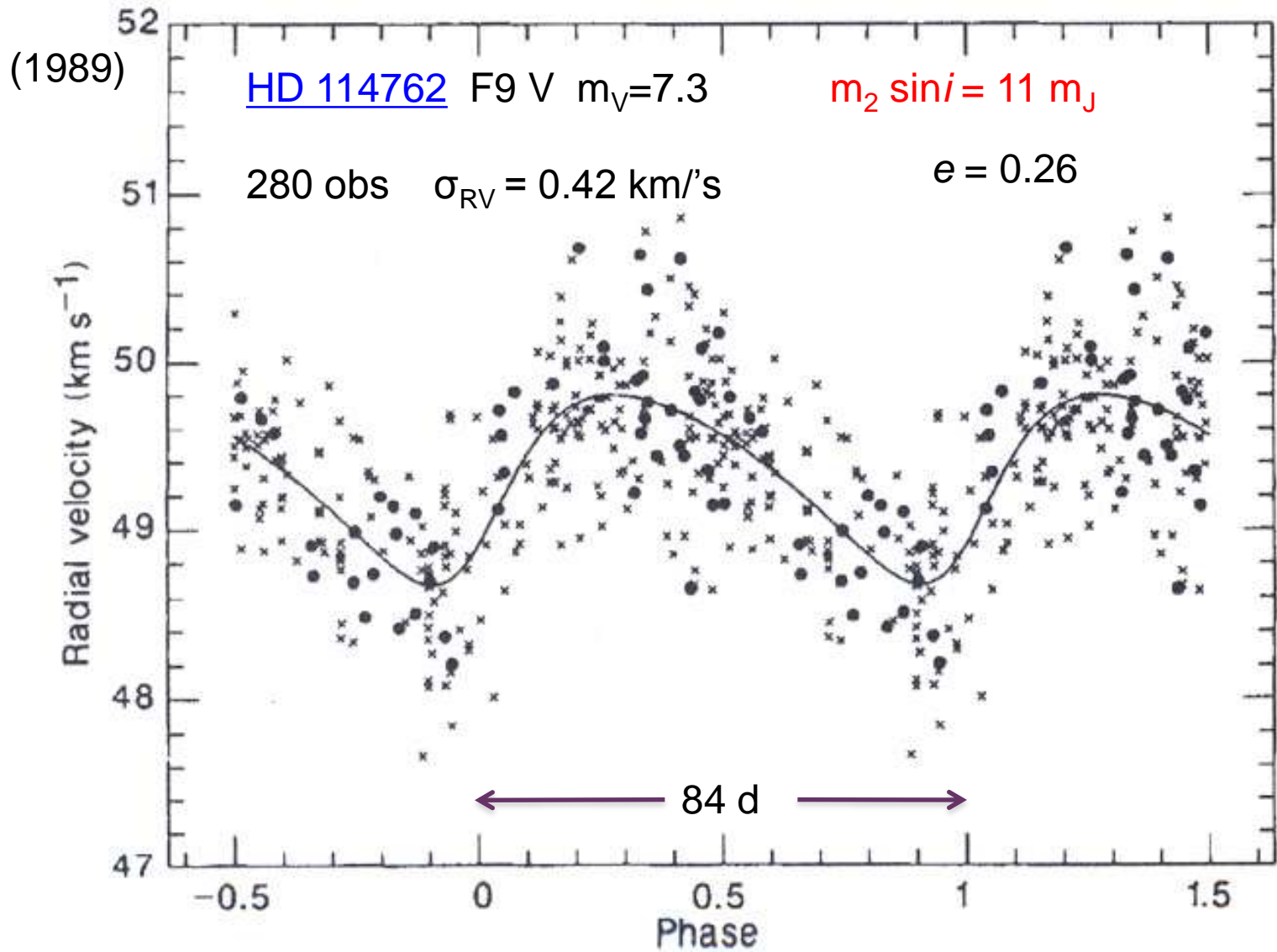
slit



RV scanner mask



Latham, Stefanik, Mazeh, Mayor, Burki -- CfA speedometer & CORAVEL



ALGOL

John Glaspey & Greg Fahlman 1973



Si diode
Vidicon
output

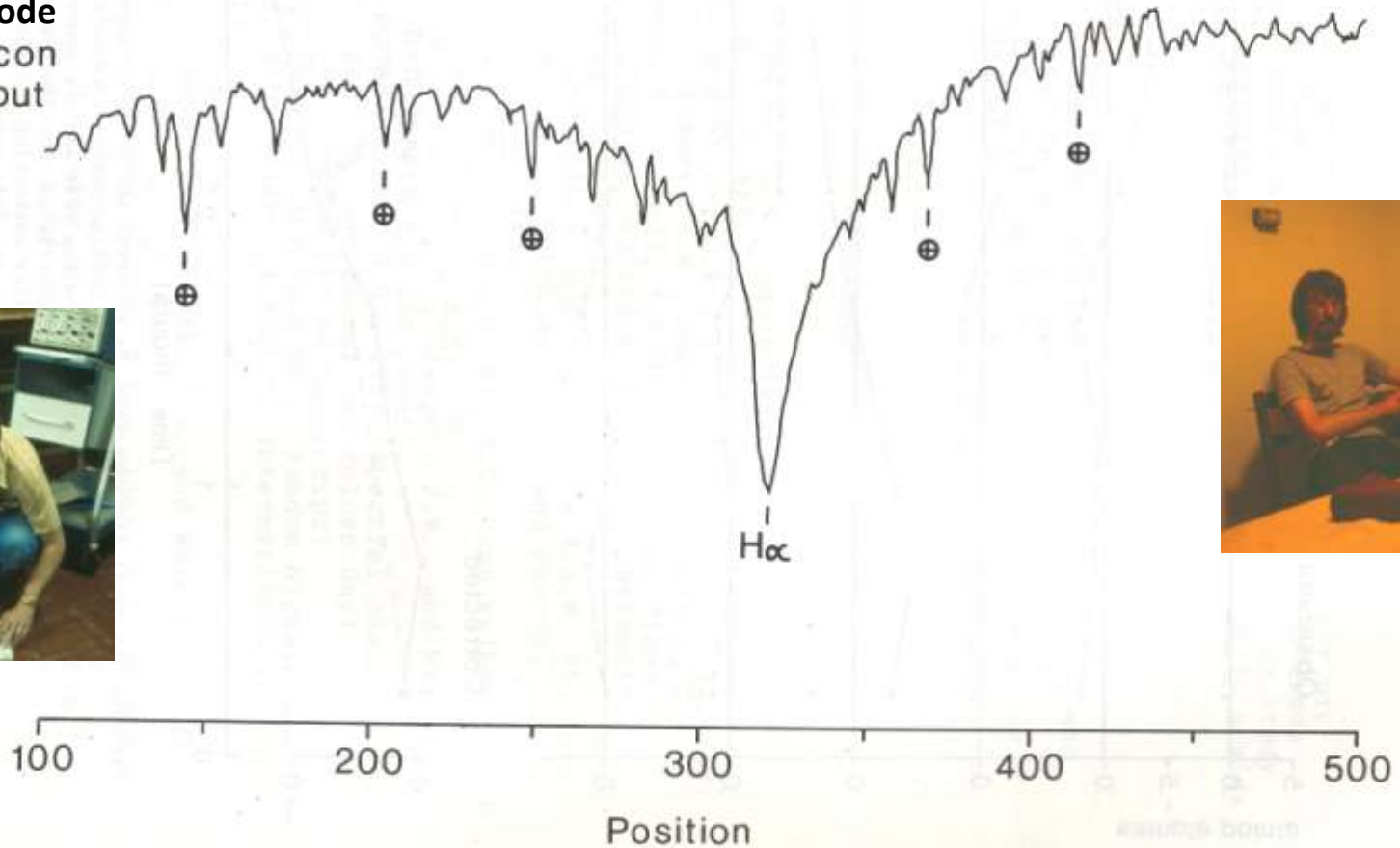


Fig. 9. Algol, Sept. 2, 1972. The five telluric features indicated were used in lieu of fiducial markers in order to determine the jitter in the Silicon Vidicon scanning raster.

ALGOL

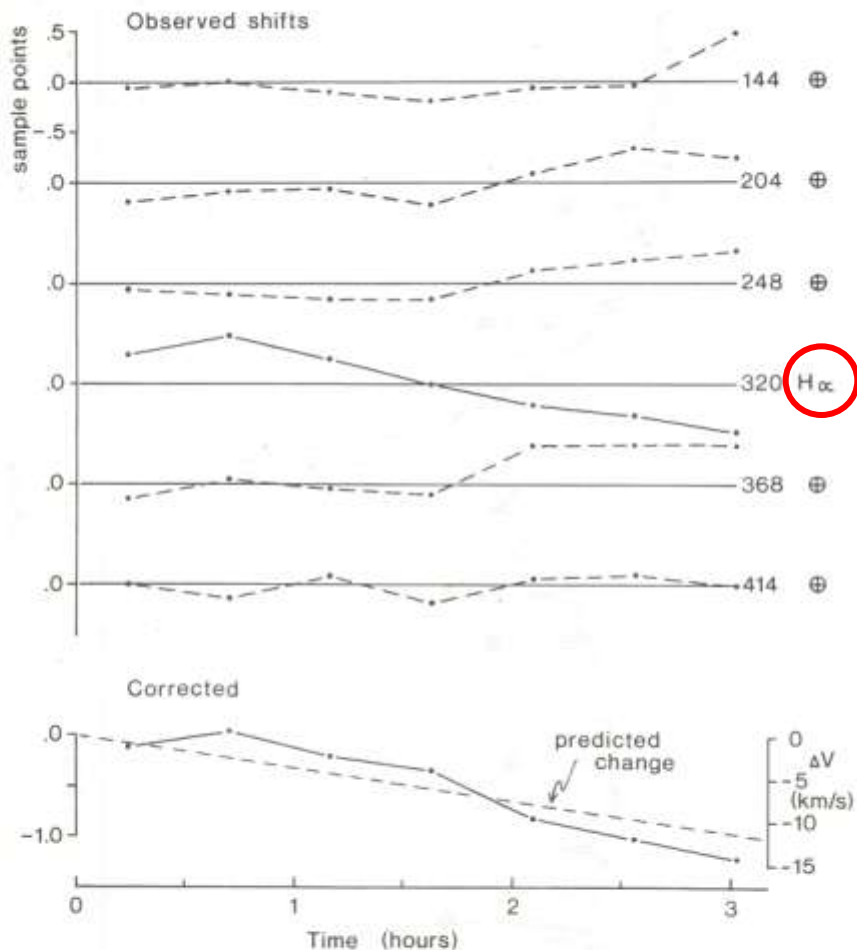


Fig. 10. The upper tracings show the observed shifts as a function of time for the H α and the five telluric lines noted on Fig. 9. The shift scale is given at the top left and the central position of each feature is listed on the right. The lower tracing shows the positional change of H α after correcting for the scanning raster variation. The scale on the left refers to the shift in sample points, and the corresponding change in radial velocity is indicated on the right. The dashed line is the change in the radial velocity of Algol A as predicted by the orbital elements of Hill et al (1971).

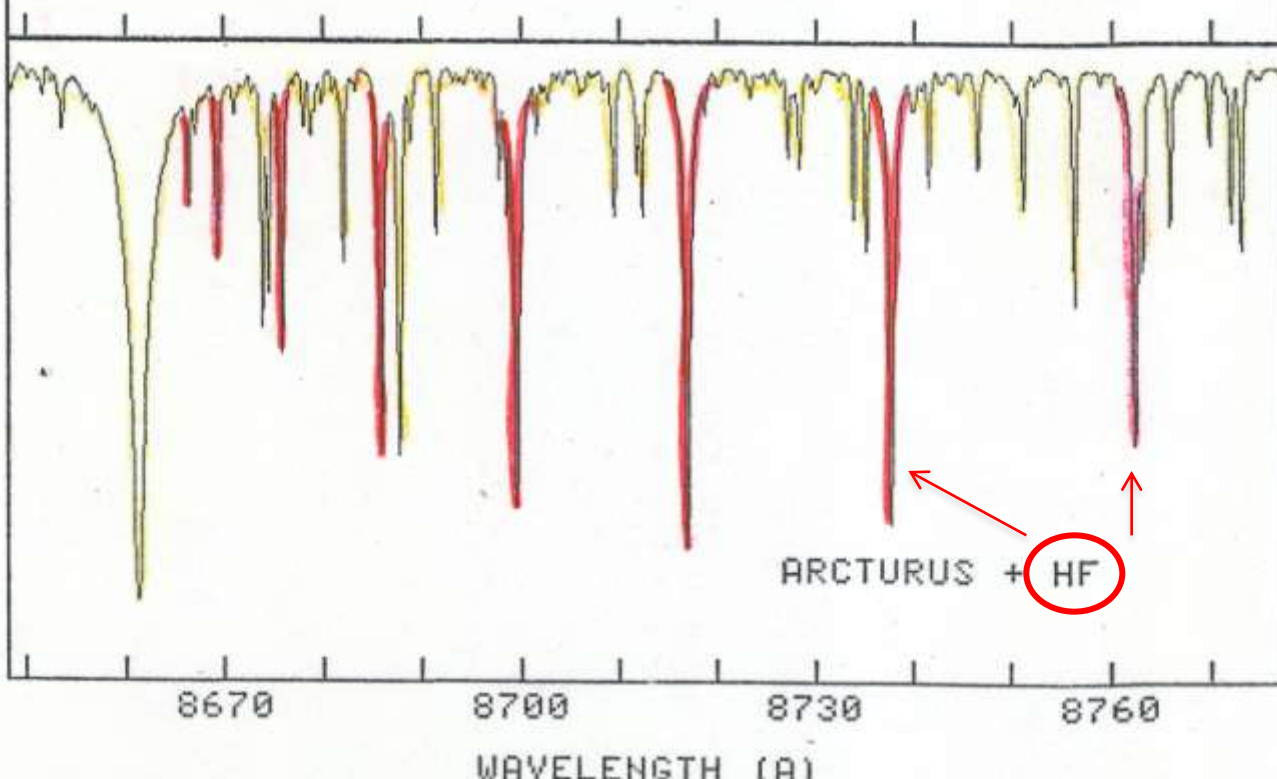
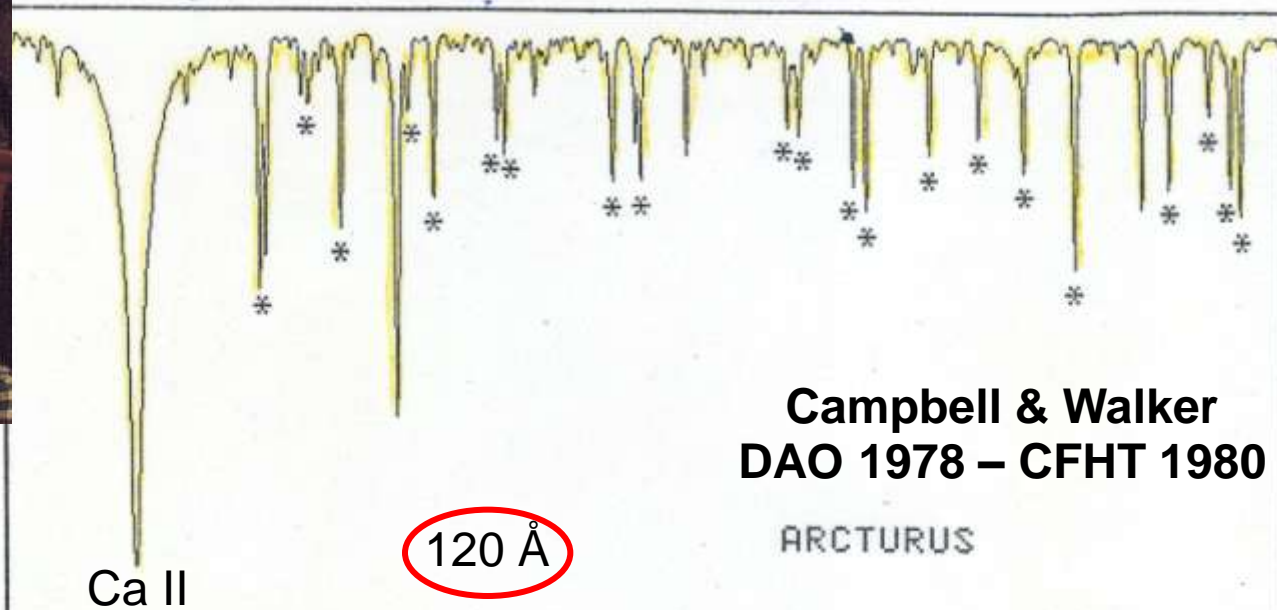
Reticon self-scanned Silicon diode arrays

256

1024

1872 15 x 750 μm



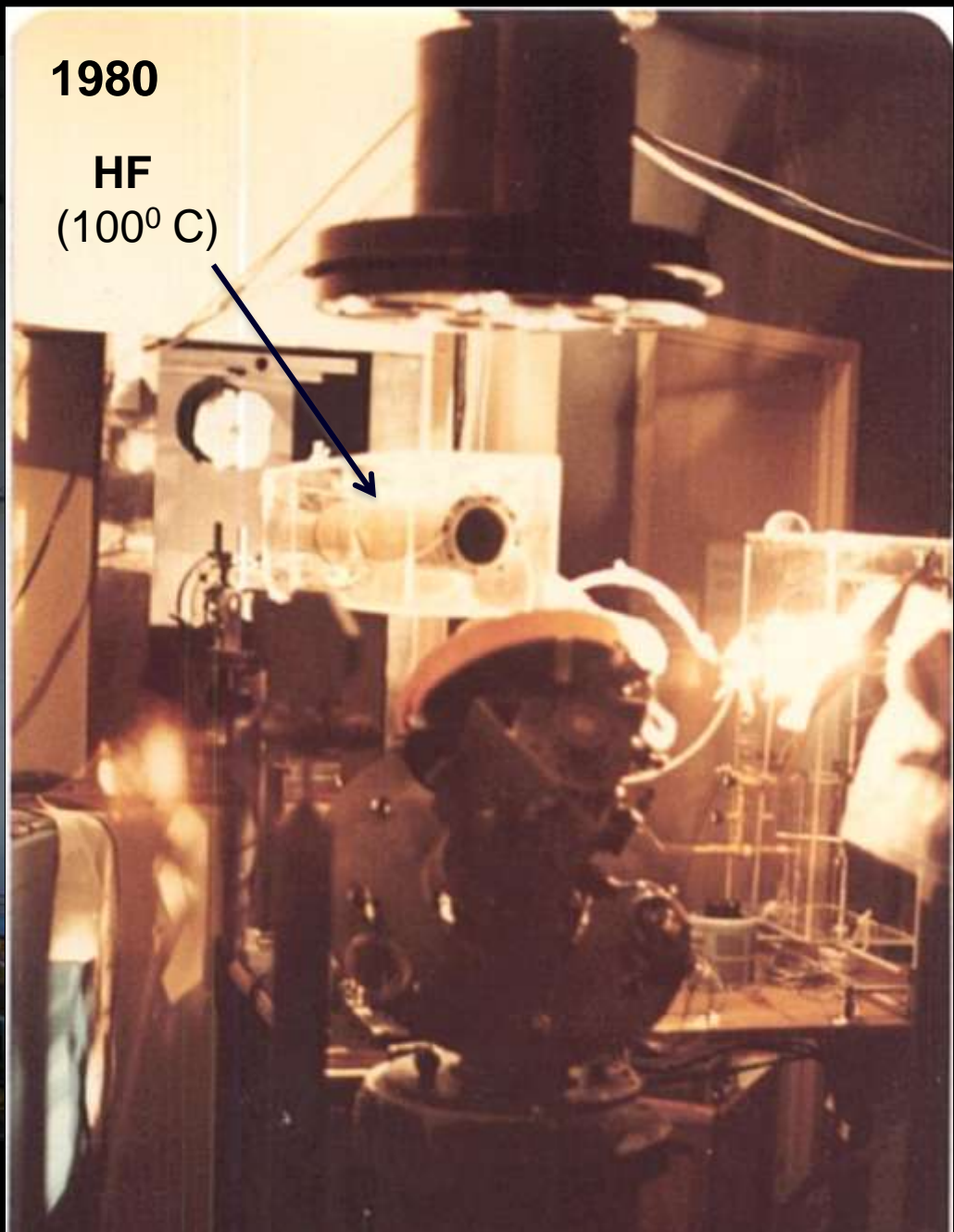


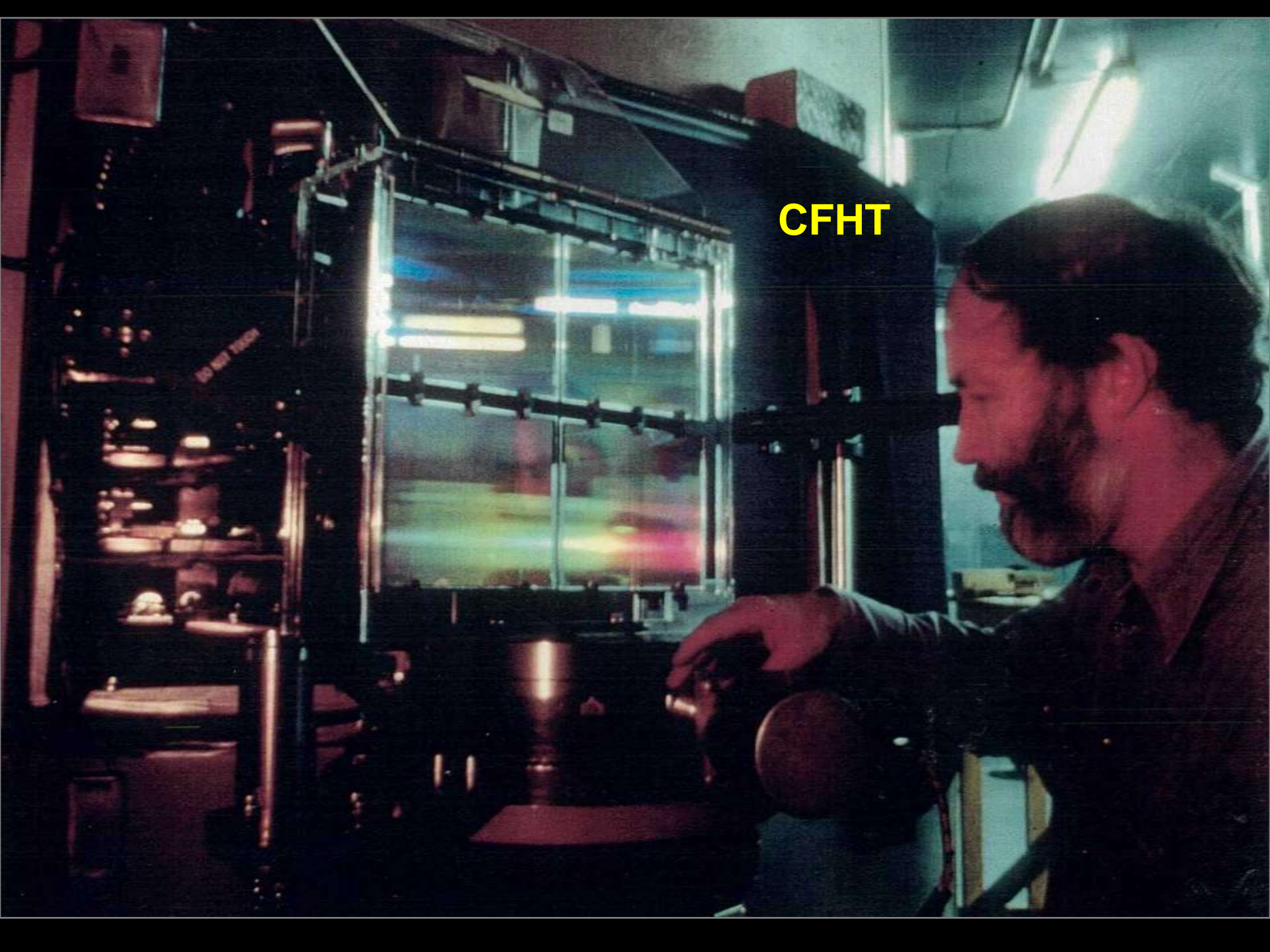
CFHT 3.6-m



1980

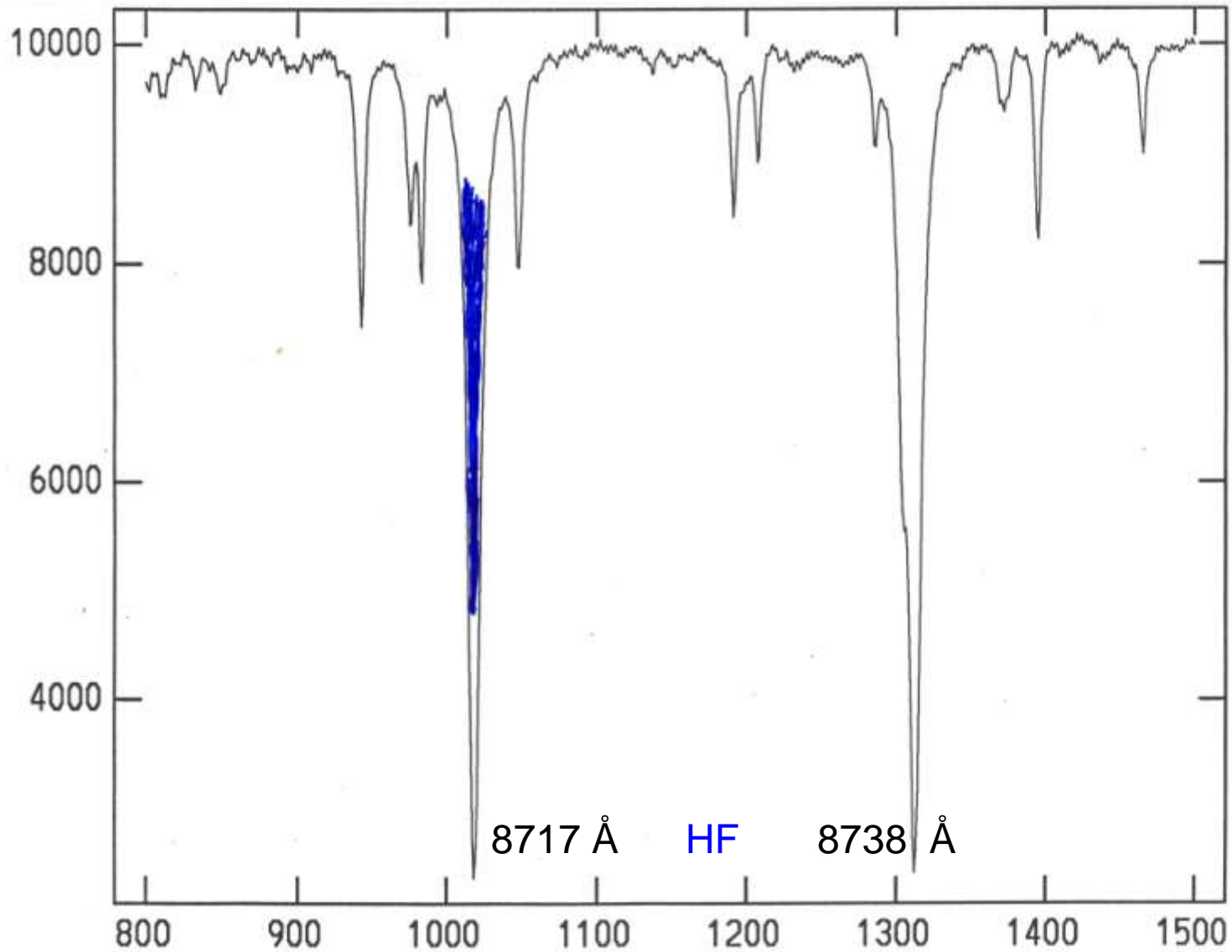
HF
(100° C)



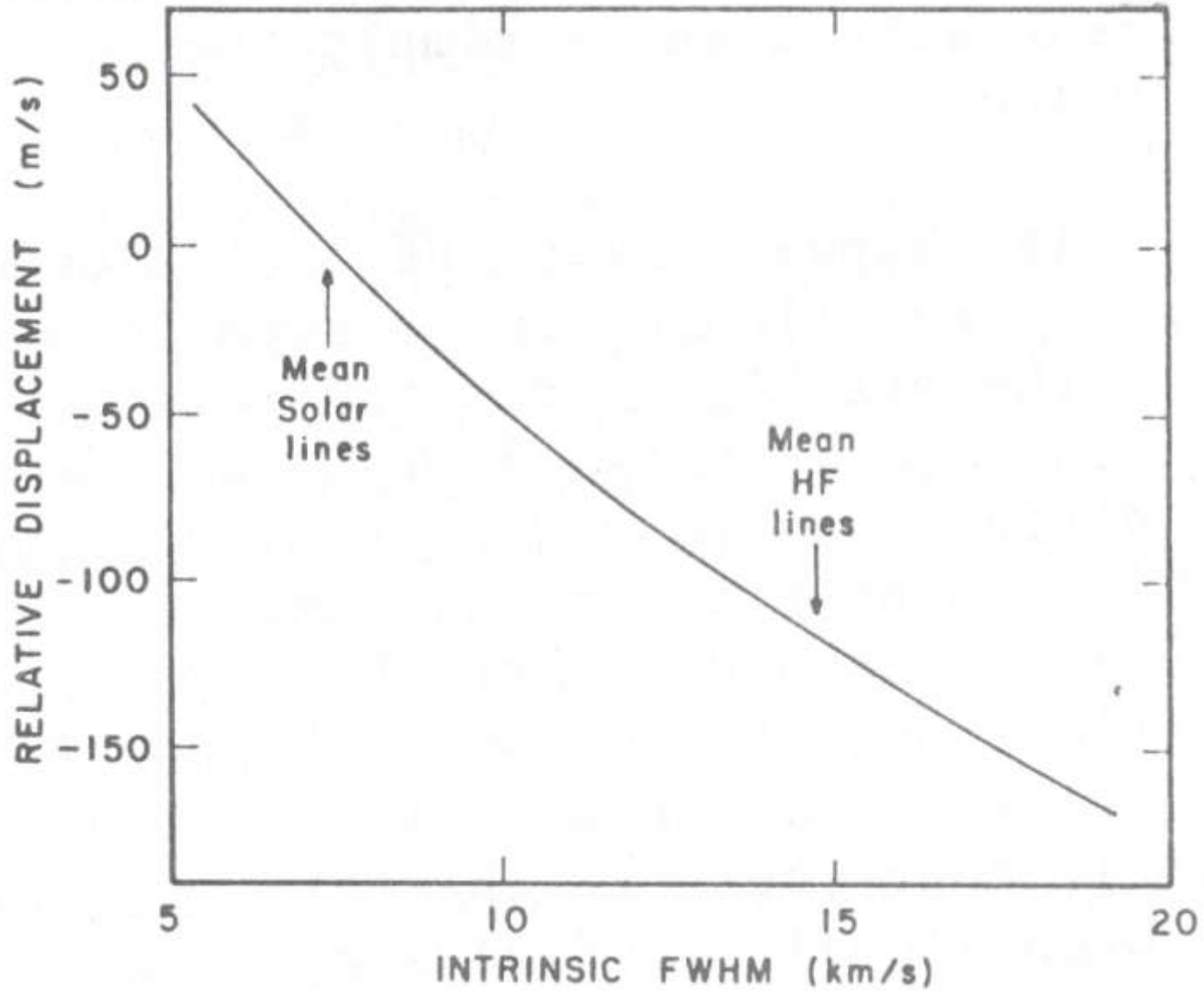


CFHT

impact of instrumental profile asymmetry ?



relative displacement of lines convolved with
a 610 m/s asymmetric instrumental profile



1988

A SEARCH FOR SUBSTELLAR COMPANIONS TO SOLAR-TYPE STARS

Campbell, Walker & Yang

12 late-type stars – 8 (4x2) nights per year [later only 6]

external error 13 m/s

flagged potential planetary companions to:

γ Cep - single line SB + superimposed 2.7 y 25 m/s amplitude

ϵ Eri

1989

YELLOW GIANTS – A NEW CLASS OF RADIAL VELOCITY VARIABLE ?

Walker, Campbell, Yang & Irwin

1992

Alex Wolszczan & Dale Frail – terrestrial companions to B1257 +12

1995

Ana Larson – β Gem 590 d planetary period [also Cochran & Hatzes]

1995

A SEARCH FOR JUPITER MASS COMPANIONS TO NEARBY STARS

21 stars

1980 – 92

Marcy & Butler

5019 to 5872 Å = 853 Å

1987



50° C

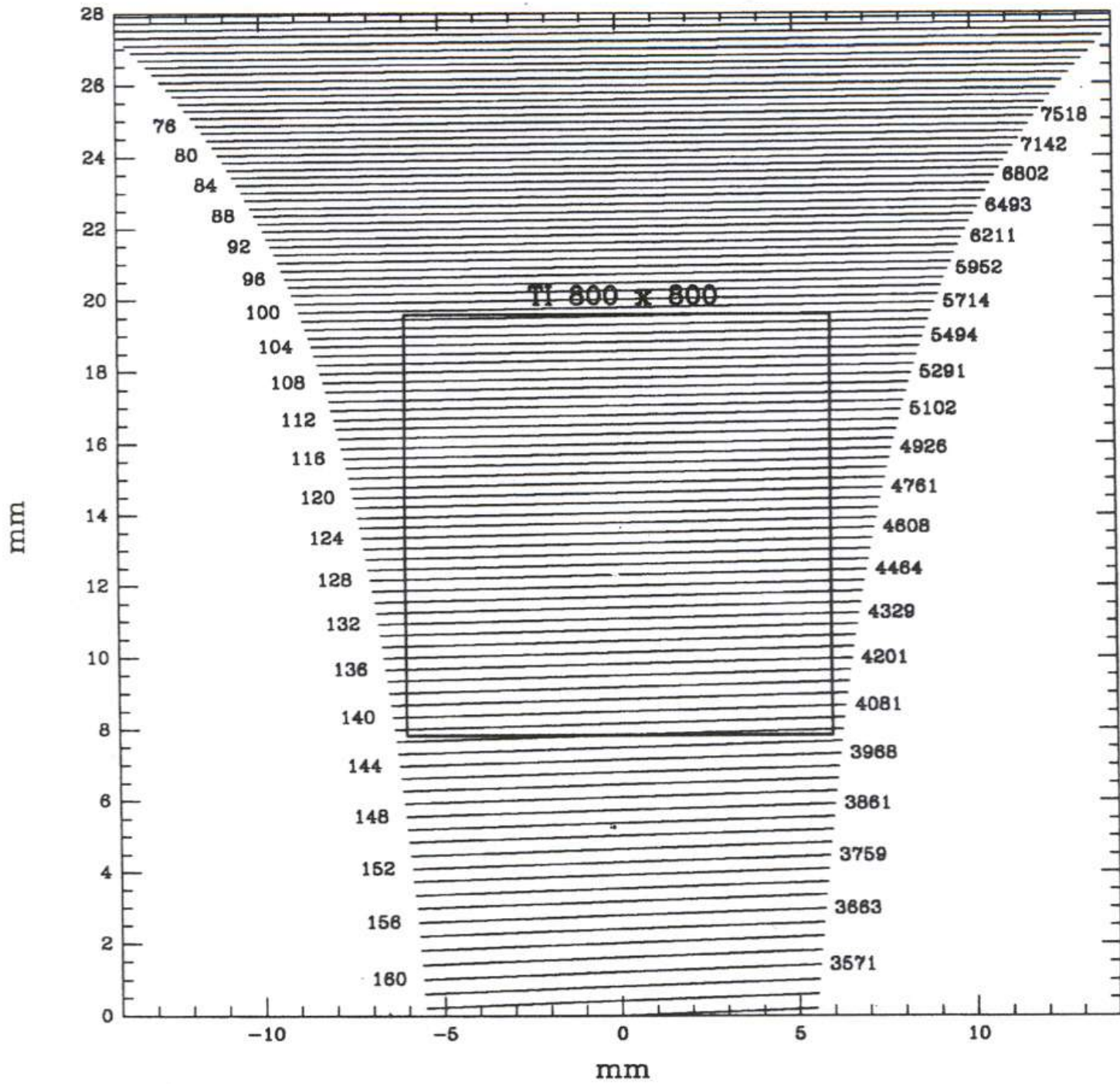
I₂

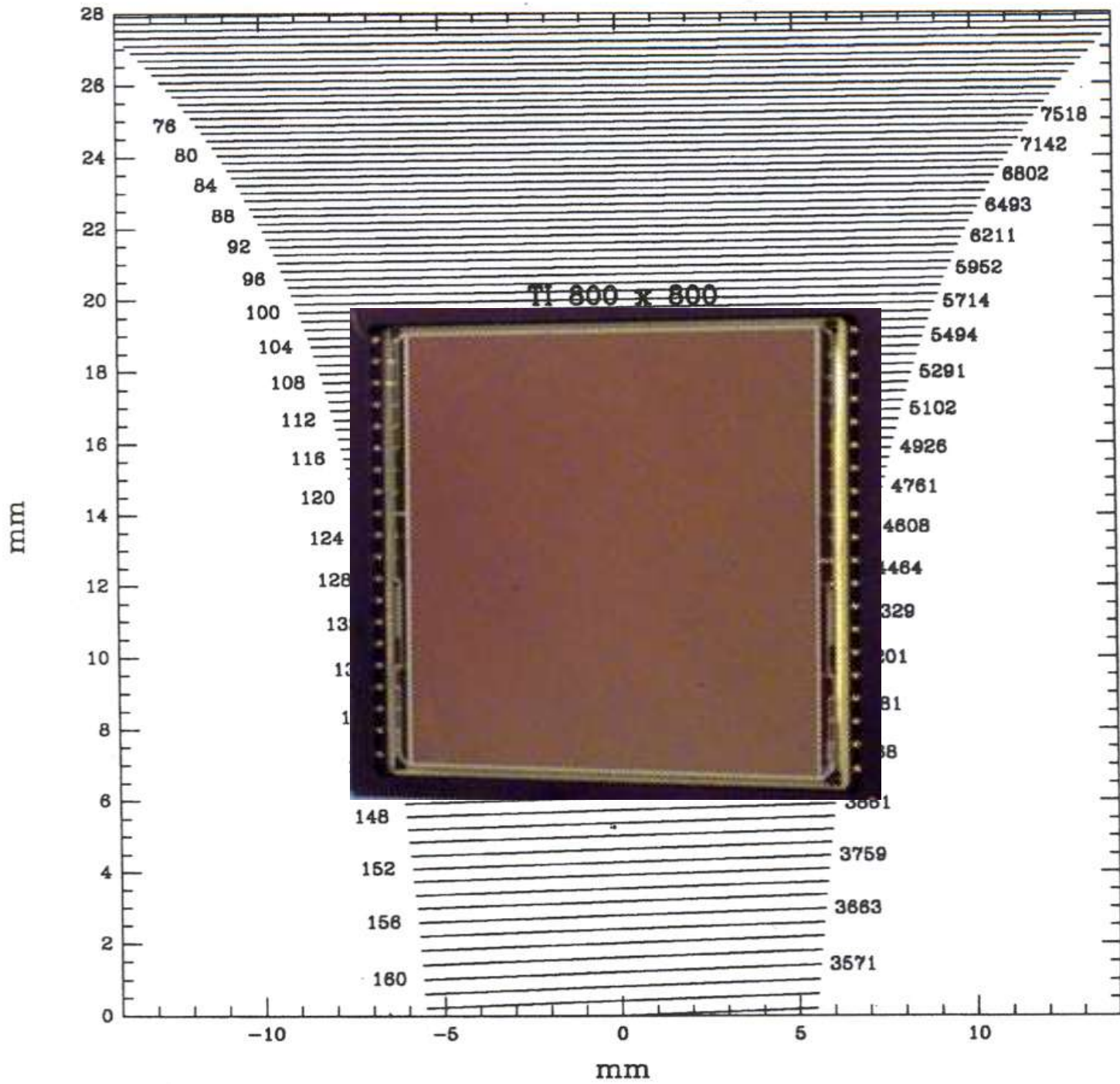
Hamilton échelle (Vogt) $R = 62,000$

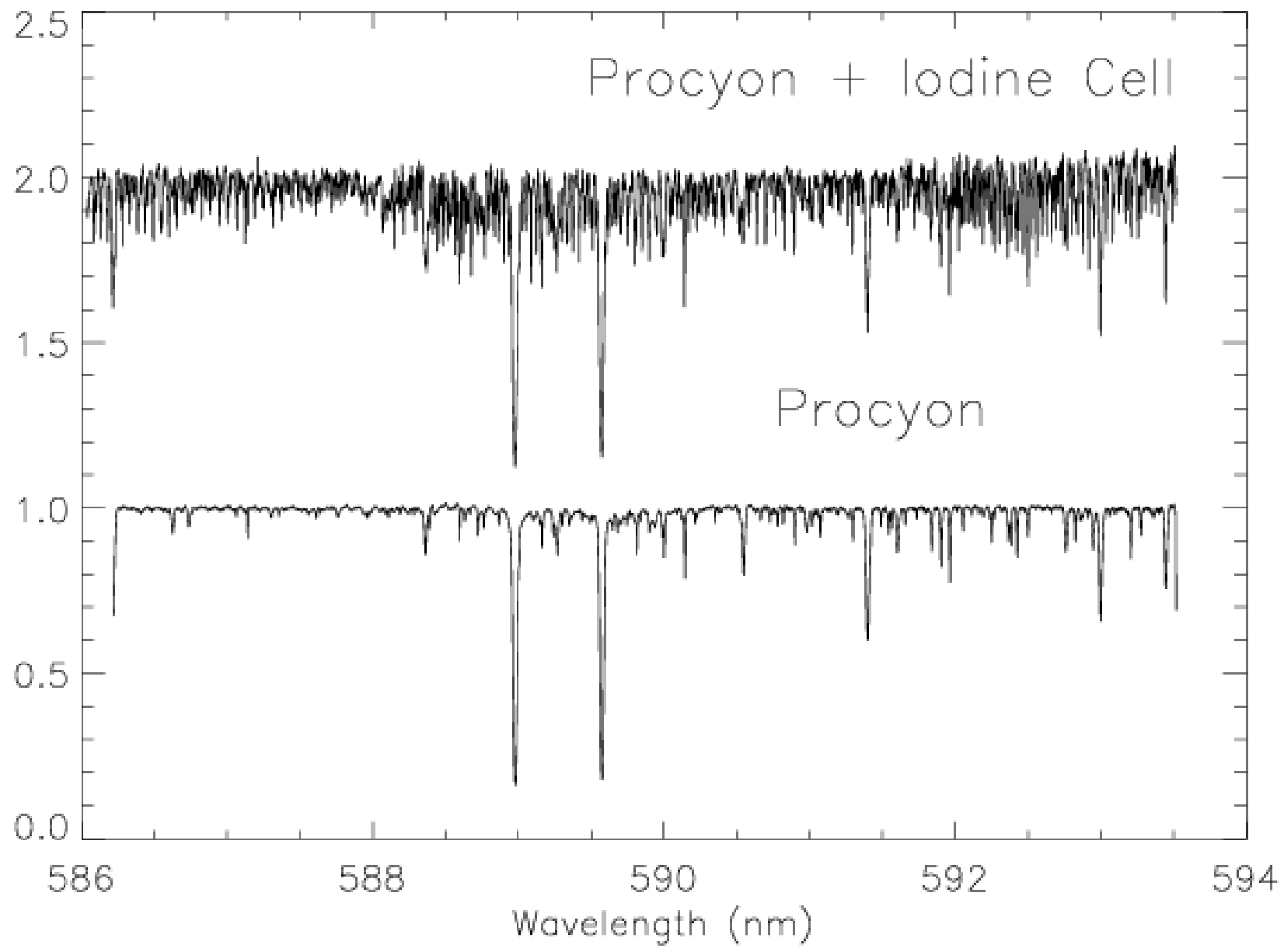


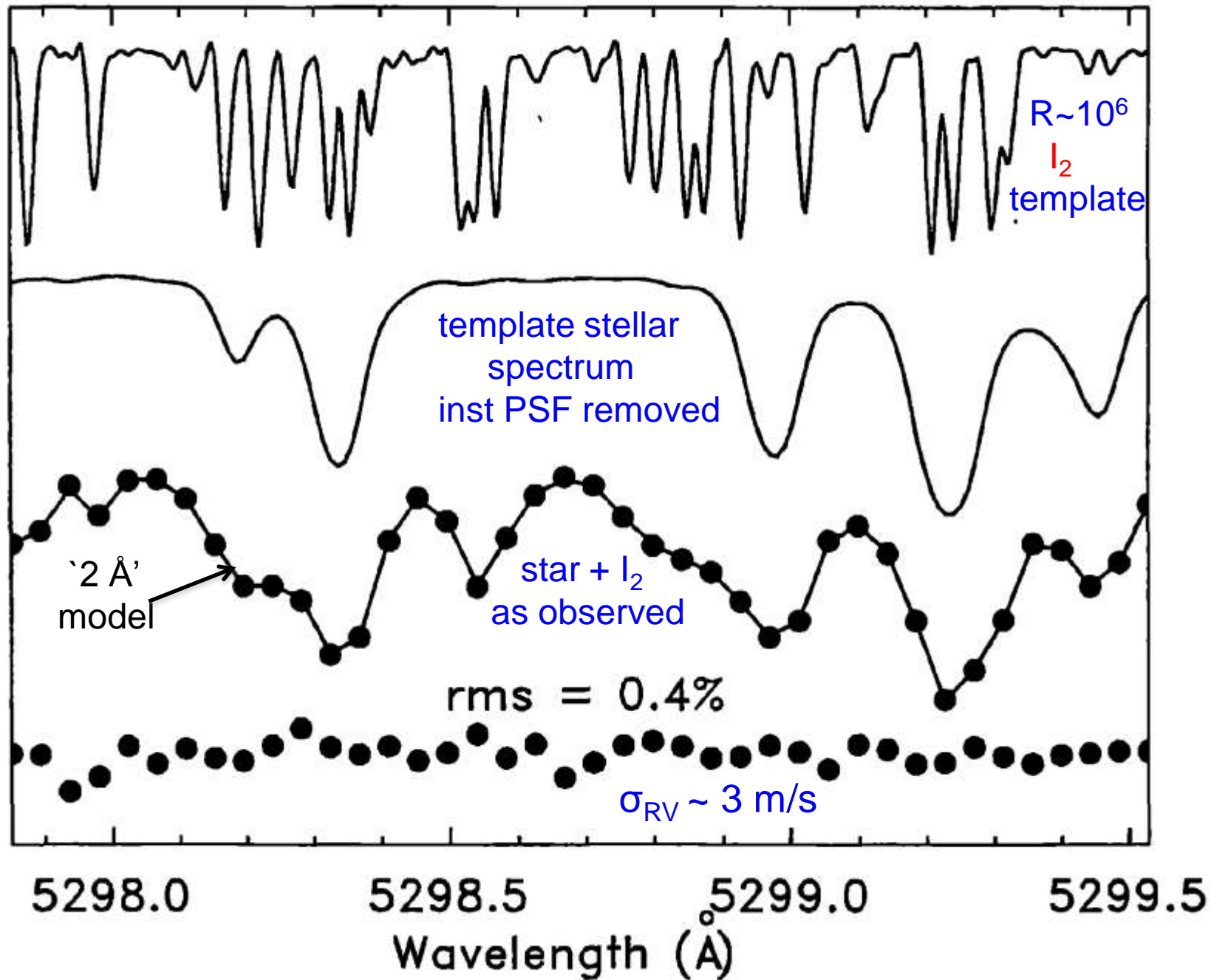
Shane
3-m





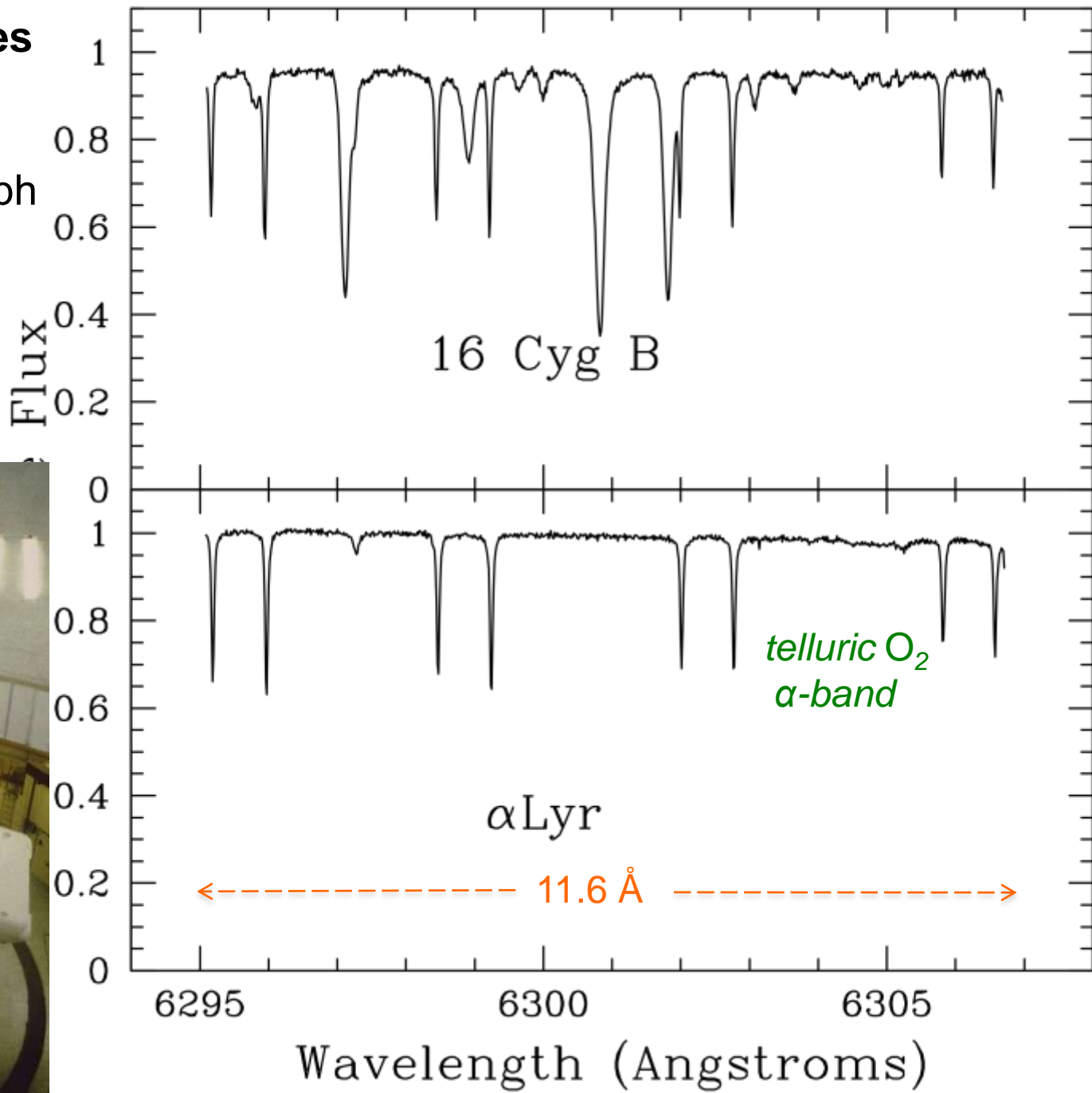






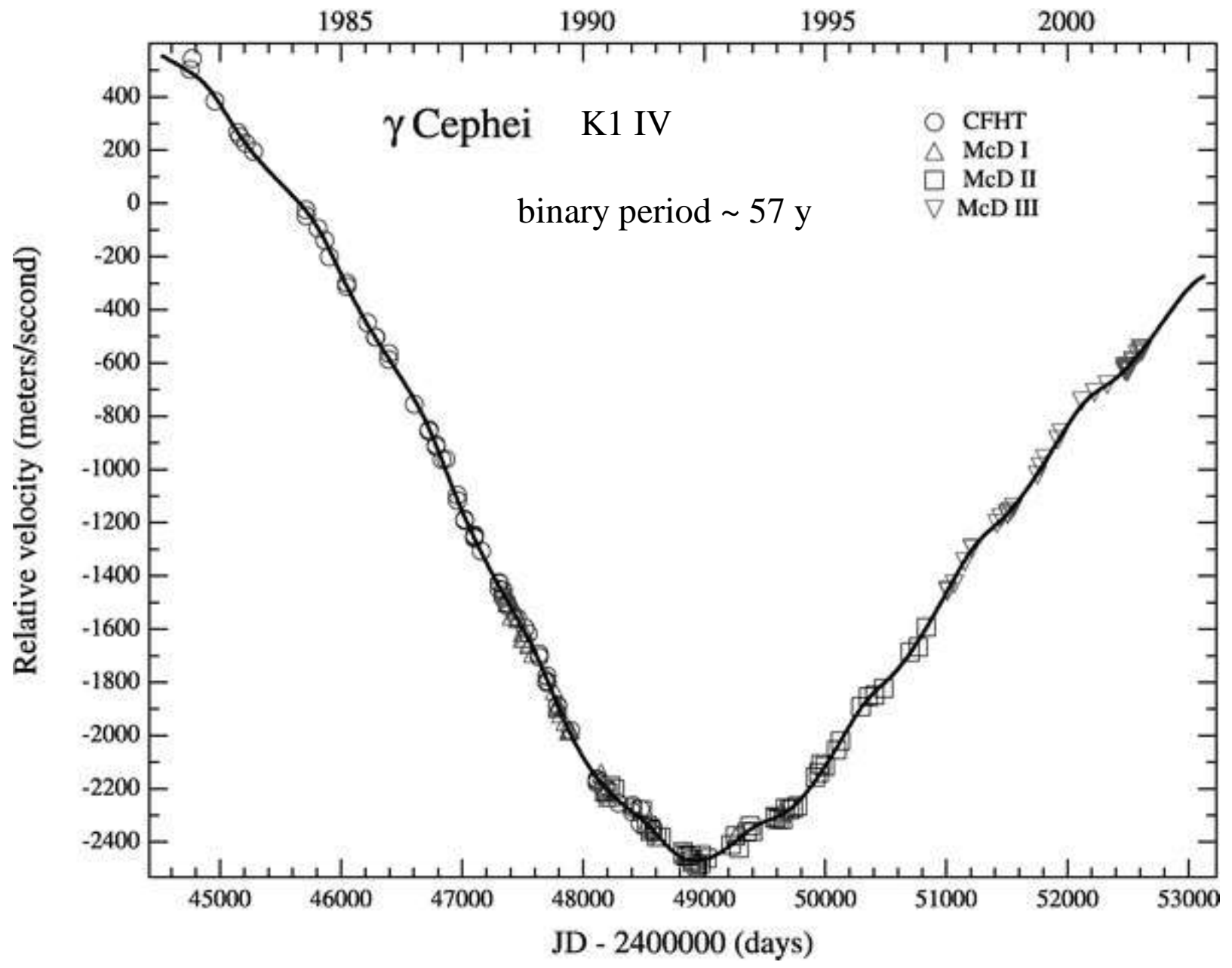
Cochran & Hatzes
MD I 1988

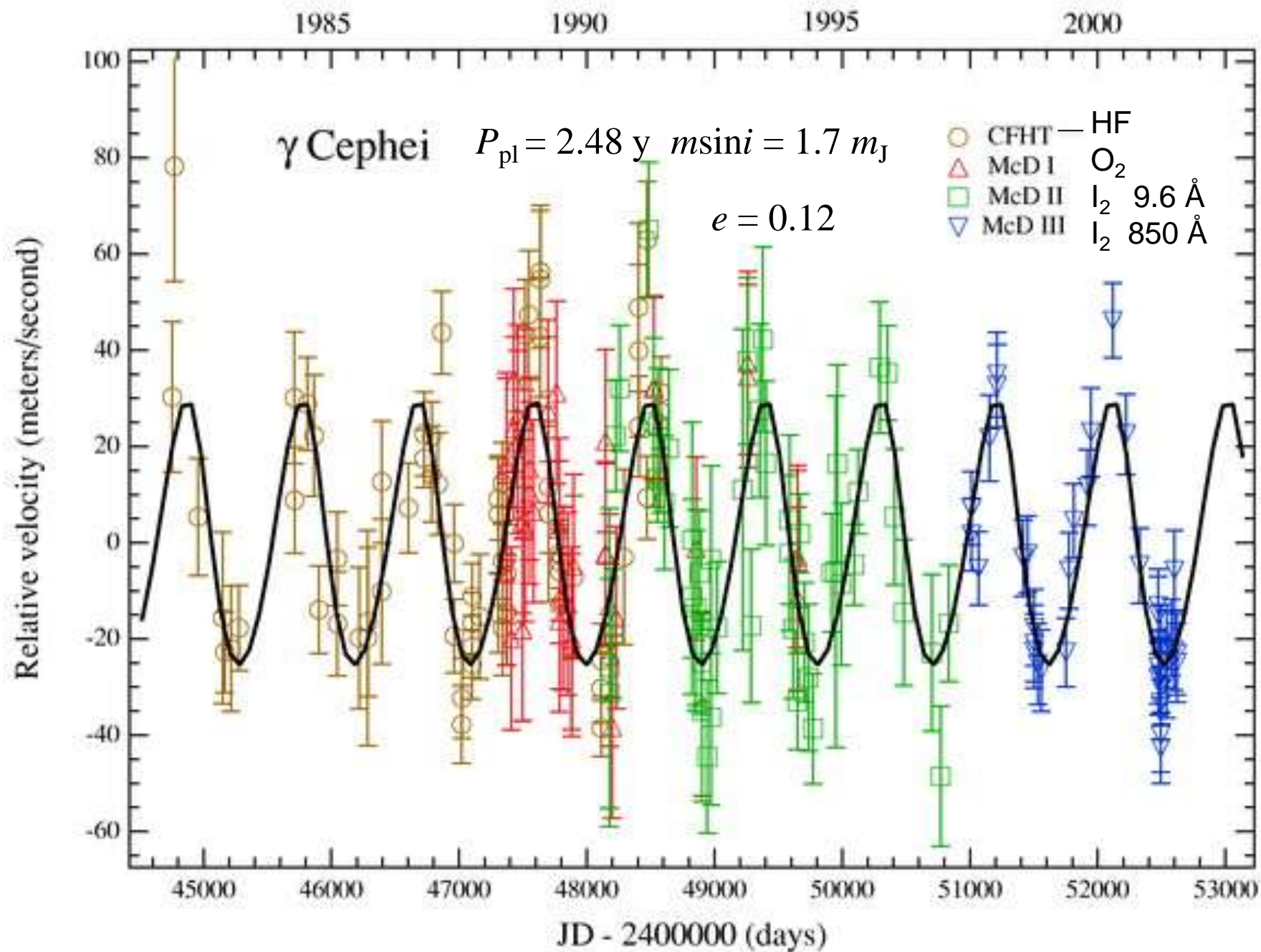
Tull coudé sp'graph
R~210,000
TI 800x800



2.7-m Harlan Smith





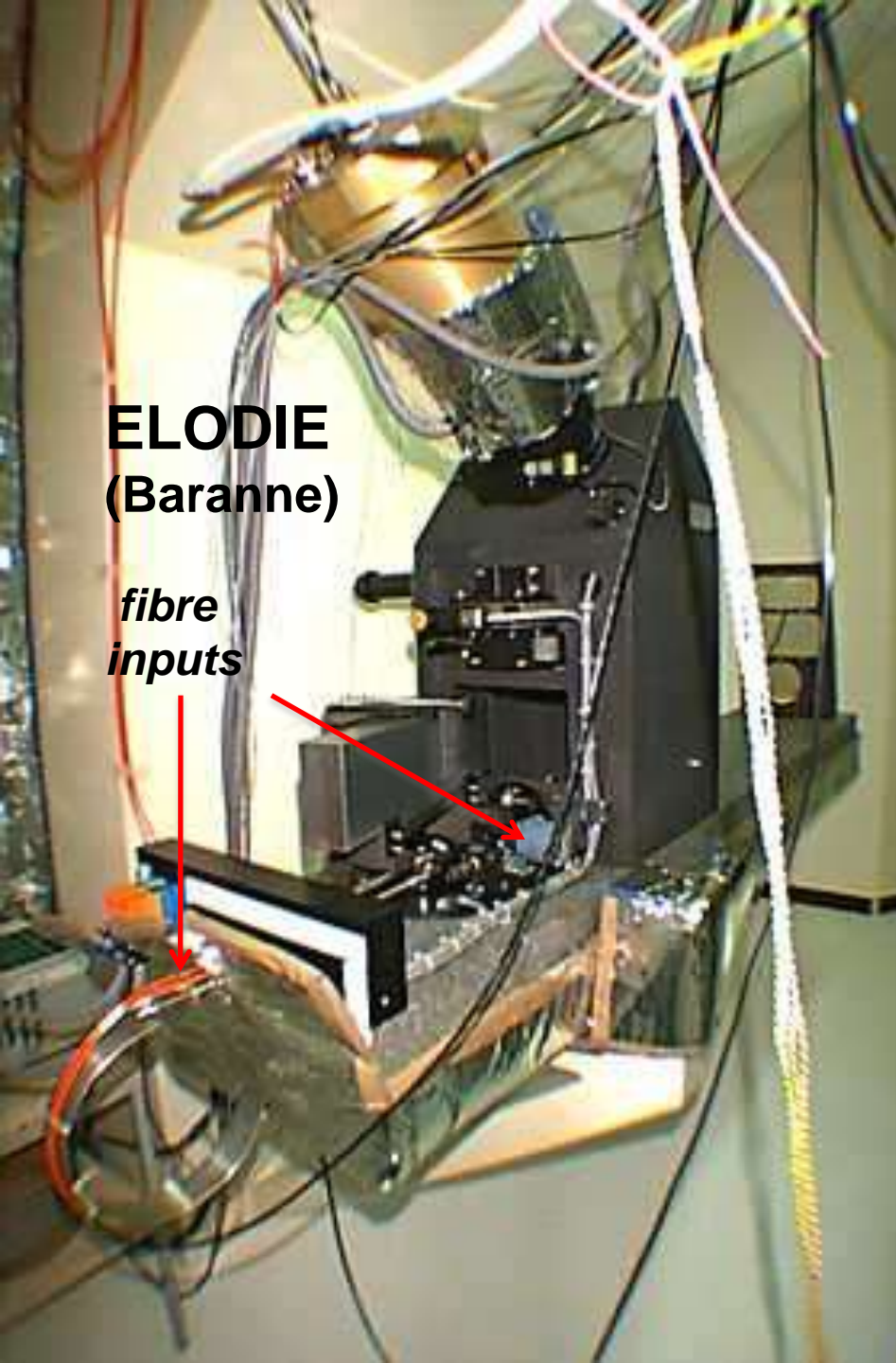


Mayor & Queloz 1994

OHP 1.93-m

ELODIE
(Baranne)

fibre
inputs



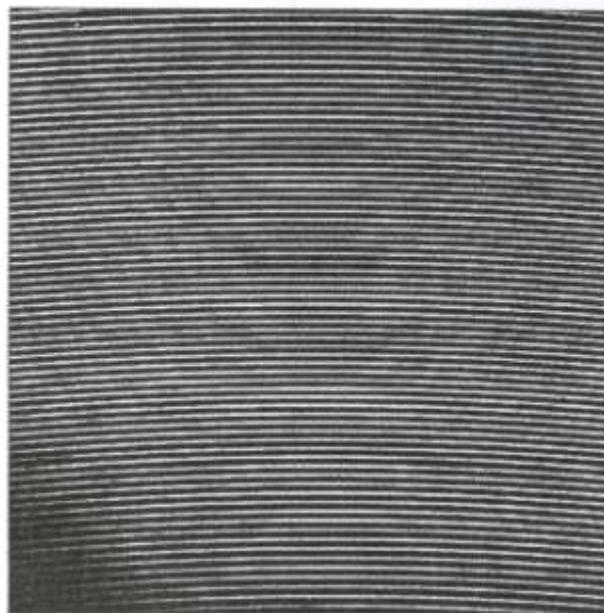
ELODIE

R = 42,000

67 orders

Tk 1024²

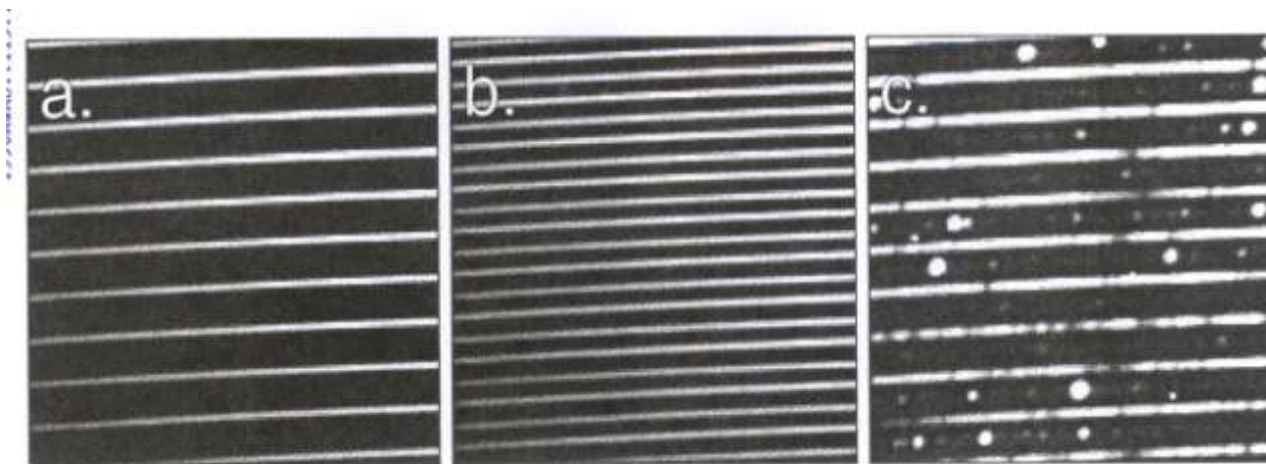
3900 to 6800 Å



*twin fibre feed
light scrambling*

→ **stable psf**

□ only



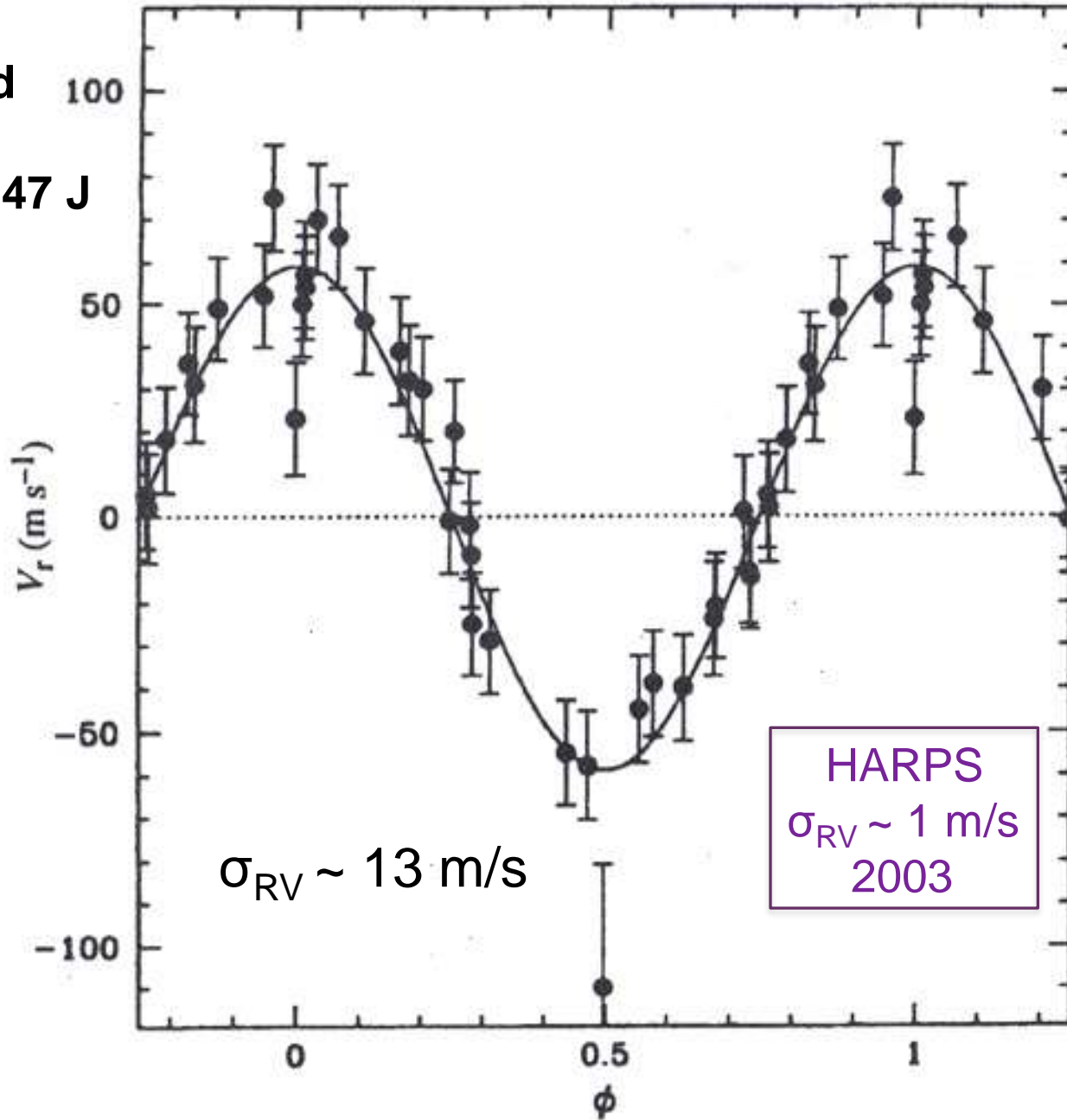
★
+ Th/Ar arc

51 Peg Mayor & Queloz

1995

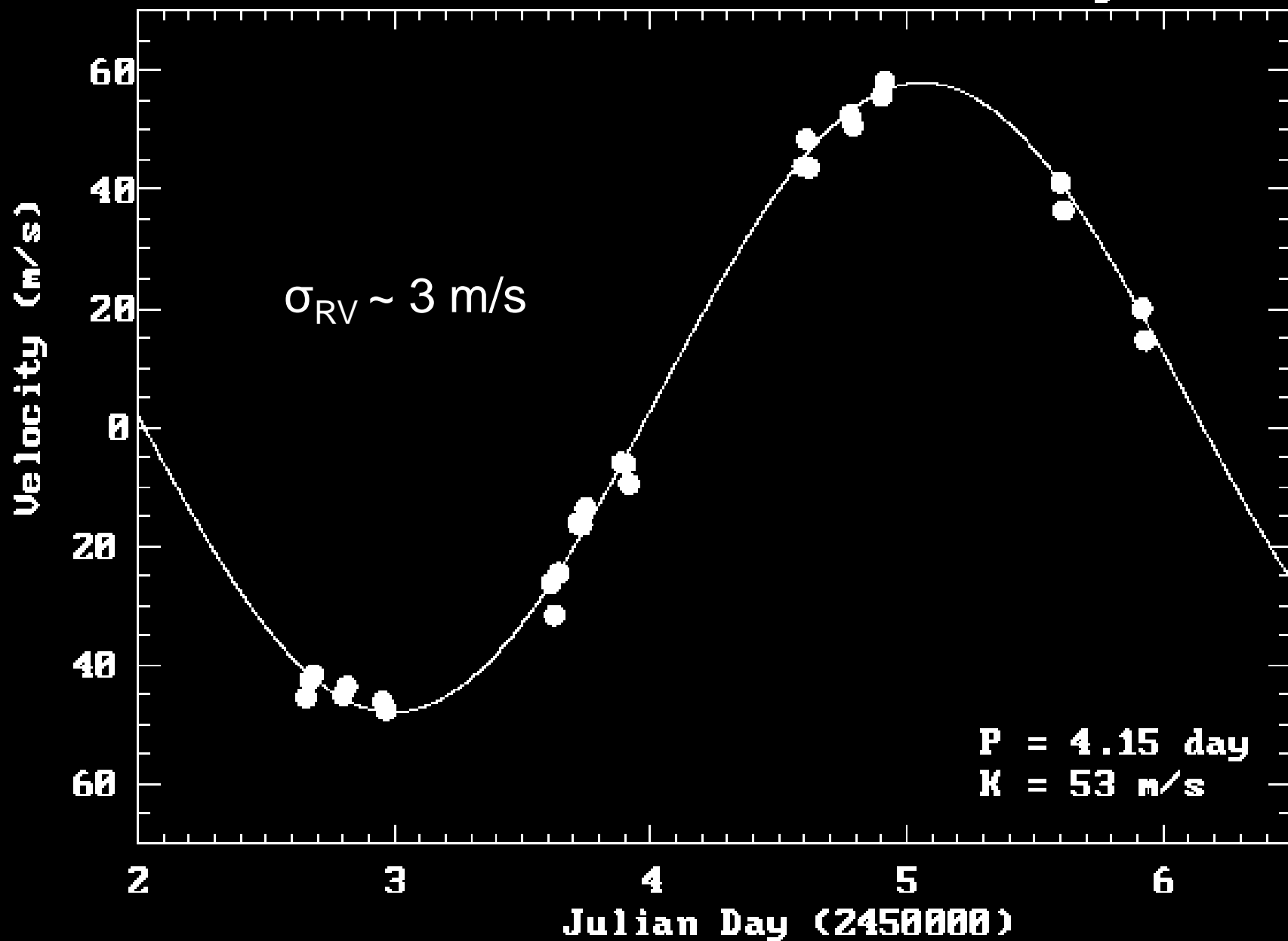
$P = 4.23$ d

$m \sin i = 0.47$ J



51 Peg

Marcy & Butler



the future ?

space

4-m at L2 – Strehl ~ 0.5

spectrograph

$R > 250,000$

λ 0.3 to 2 μm